

CS-184: Computer Graphics

Lecture 17: Radiometry

Maneesh Agrawala
University of California, Berkeley

Today

Radiometry: measuring light

- Local Illumination and Raytracing were discussed in an *ad hoc* fashion
- Proper discussion requires proper units
- Not just pretty pictures... but correct pictures

2

Matching Reality



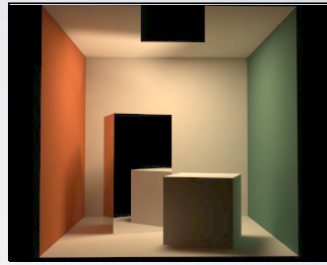
Unknown

3

Matching Reality



Photo



Rendered



Cornell Box Comparison
Cornell Program of Computer Graphics

4

Units

Light energy

- Really power not energy is what we measure
- Joules / second (J/s) = Watts (W)

Spectral energy density

- Power per unit spectrum interval
- Watts / nano-meter (W/nm)
- Properly done as function over spectrum
- Often just sampled for RGB

Often we assume people know we're talking about S.E.D. and just say E...

5

Irradiance

Total light striking surface from all directions

- Only meaningful w.r.t. a surface
- Power per square meter ($\mathbf{W/m^2}$)
- Really S.E.D. per square meter ($\mathbf{W/m^2 /nm}$)
- Not all directions sum the same because of foreshortening



6

Radiant Exitance

Total light *leaving* surface over all directions

- Only meaningful w.r.t. a surface
- Power per square meter (W/m^2)
- Really S.E.D. per square meter ($\text{W}/\text{m}^2 / \text{nm}$)
- Also called Radiosity
- Sum over all directions \Rightarrow same in all directions



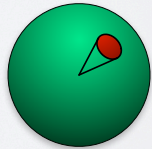
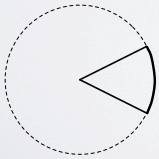
Solid Angles

Regular angles measured in *radians* $[0..2\pi]$

- Measured by arc-length on unit circle

Solid angles measured in *steradians* $[0..4\pi]$

- Measured by area on unit sphere
- Not necessarily little round pieces...



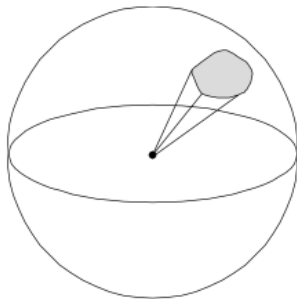
Angles and Solid Angles

■ **Angle** $\theta = \frac{l}{r}$

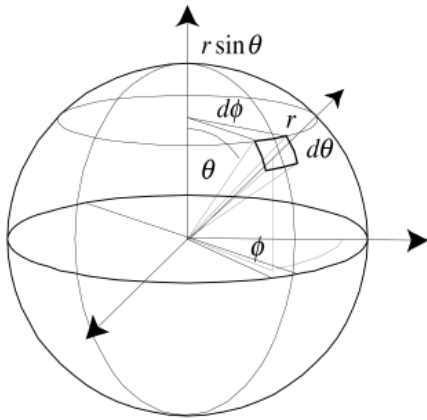
\Rightarrow circle has 2π radians

■ **Solid angle** $\Omega = \frac{A}{R^2}$

\Rightarrow sphere has 4π steradians



Differential Solid Angles



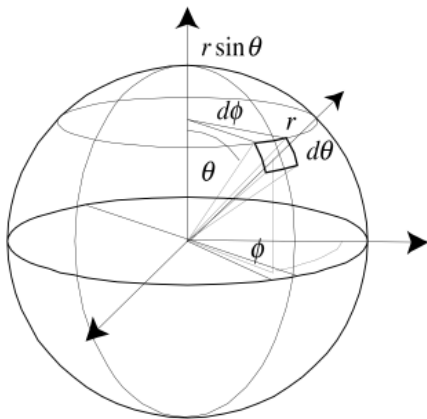
$$dA = (r d\theta)(r \sin \theta d\phi)$$

$$= r^2 \sin \theta d\theta d\phi$$

CS348B Lecture 4

Pat Hanrahan, 2009

Differential Solid Angles



$$dA = (r d\theta)(r \sin \theta d\phi)$$

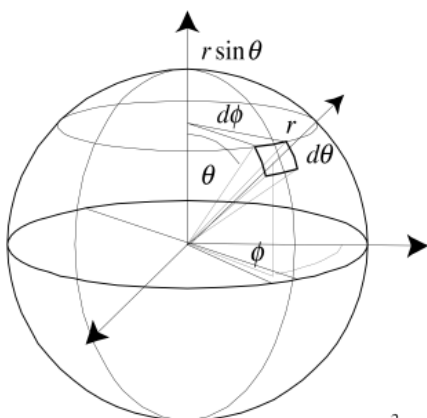
$$= r^2 \sin \theta d\theta d\phi$$

$$d\omega = \frac{dA}{r^2} = \sin \theta d\theta d\phi$$

CS348B Lecture 4

Pat Hanrahan, 2009

Differential Solid Angles



$$d\omega = \sin \theta d\theta d\phi$$

$$\Omega = \int_{S^2} d\omega$$

$$= \int_0^\pi \int_0^{2\pi} \sin \theta d\theta d\phi$$

$$= \int_{-1}^1 \int_0^{2\pi} d\cos \theta d\phi$$

$$= 4\pi$$

Sphere S^2

CS348B Lecture 4

Pat Hanrahan, 2009

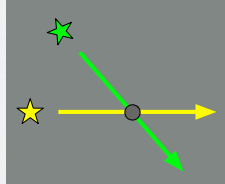
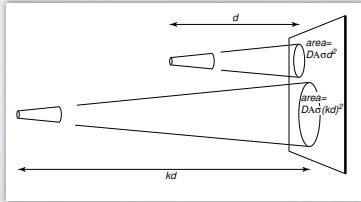
Radiance

Light energy passing through a point in space within a given solid angle

- Energy per steradian per square meter ($\text{W}/\text{m}^2 / \text{sr}$)
- S.E.D. per steradian per square meter ($\text{W}/\text{m}^2 / \text{sr} / \text{nm}$)

Constant along straight lines in free space

- Area of surface being sampled is proportional to distance and light inversely proportional to squared distance

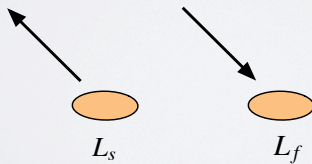


13

Radiance

Near surfaces, differentiate between

- Radiance from the surface (surface radiance)
- Radiance from other things (field radiance)



14

Light Fields

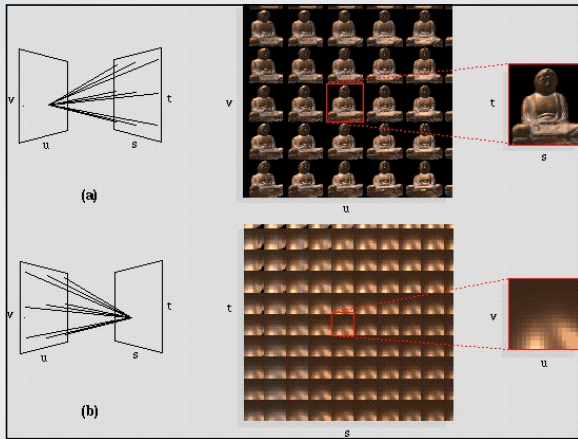
Radiance at every point in space, direction, and frequency: 6D function

Collapse frequency to RGB, and assume free space: 4D function

Sample and record it over some volume

15

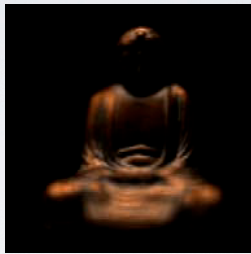
Light Fields



Levoy and Hanrahan, SIGGRAPH 1996

16

Light Fields



Levoy and Hanrahan, SIGGRAPH 1996

17

Light Fields



Michelangelo's *Statue of Night*
From the Digital Michelangelo Project

18

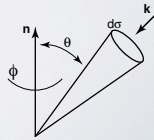
Computing Irradiance

Integrate incoming radiance (field radiance) over all direction

- Take into account foreshortening

$$H = \int_{\Omega} L_f(\mathbf{k}) \cos(\theta) d\sigma$$

$$H = \int_0^{2\pi} \int_0^{\pi/2} L_f(\theta, \phi) \cos(\theta) \sin(\theta) d\theta d\phi$$



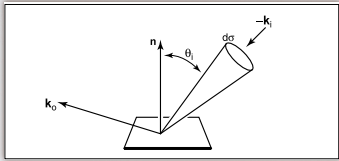
19

Revisiting The BRDF

How much light from direction \mathbf{k}_i goes out in direction \mathbf{k}_o

Now we can talk about units:

- BRDF is ratio of surface radiance to the foreshortened field radiance multiplied by small solid angle (denom. similar to integrand of irradiance - see prev. slide)



$$\rho(\mathbf{k}_i, \mathbf{k}_o) = \frac{L_s(\mathbf{k}_o)}{L_f(\mathbf{k}_i) \cos(\theta_i) \Delta\sigma_i}$$

We left out frequency dependence here...

Also note for perfect Lambertian reflector with constant BRDF $\rho = 1/\pi$

20

The Rendering Equation

Total light going out in some direction is given by an integral over all incoming directions:

$$L_s(\mathbf{k}_o) = \int_{\Omega} \rho(\mathbf{k}_i, \mathbf{k}_o) L_f(\mathbf{k}_i) \cos(\theta_i) d\sigma_i$$

- Note, this is recursive (my L_f is another's L_s)

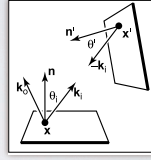
21

The Rendering Equation

$$L_s(\mathbf{k}_o) = \int_{\Omega} \rho(\mathbf{k}_i, \mathbf{k}_o) L_f(\mathbf{k}_i) \cos(\theta_i) d\sigma_i$$

Rewrite explicitly in terms of surface radiances only

$$L_f(\mathbf{k}_i) = L_s(-\mathbf{k}_i) \quad \Delta\sigma_i = \frac{\Delta A' \cos(\theta')}{\|\mathbf{x} - \mathbf{x}'\|^2}$$



$$L_s(\mathbf{x}, \mathbf{k}_o) = \int_{x' \text{ visible to } x} \frac{\rho(\mathbf{k}_i, \mathbf{k}_o) L_s(\mathbf{x}', \mathbf{x} - \mathbf{x}') \cos(\theta_i) \cos(\theta')}{\|\mathbf{x} - \mathbf{x}'\|^2} dA'$$

$$L_s(\mathbf{x}, \mathbf{k}_o) = \int_{\text{all } x'} \frac{\rho(\mathbf{k}_i, \mathbf{k}_o) L_s(\mathbf{x}', \mathbf{x} - \mathbf{x}') \delta(\mathbf{x}, \mathbf{x}') \cos(\theta_i) \cos(\theta')}{\|\mathbf{x} - \mathbf{x}'\|^2} dA'$$

$$\delta(\mathbf{x}, \mathbf{x}') = \begin{cases} 1 & \text{if } \mathbf{x} \text{ and } \mathbf{x}' \text{ are mutually visible} \\ 0 & \text{otherwise} \end{cases}$$

22

Light Paths

Many paths from light to eye

Characterize by the types of bounces

- Begin at light
- End at eye
- "Specular" bounces
- "Diffuse" bounces



23

Light Paths

Describe paths using strings

- LDE, LDSE, LSE, etc.

Describe types of paths with regular expressions

- $L\{D\}S^*E$ ← Visible paths
- $L\{D\}S^*S^*E$ ← Standard raytracing
- $L\{D\}S\}E$ ← Local illumination
- LD^*E ← Radiosity method
(have not talked about yet)

24